

$U(1)_\chi$ and Seesaw Dirac Neutrinos

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Abstract

In the context of $SO(10) \rightarrow SU(5) \times U(1)_\chi$, it is shown how seesaw Dirac neutrinos may be obtained. In this framework, $U(1)$ lepton number is conserved, with which self-interacting dark matter with a light scalar dilepton mediator may be implemented. In addition, $U(1)$ baryon number may be broken to $(-1)^{3B}$, thereby generating a baryon asymmetry of the Universe. The axionic solution to the strong CP problem may also be incorporated.

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1. INTRODUCTION

In considering $SO(10)$ grand unification, the common approach is to allow an intermediate step with left-right symmetry, i.e. $SU(4) \times SU(2)_L \times SU(2)_R$ or $SU(3)_C \times SU(2)_L \times SU(2)_R \times U(1)_{(B-L)/2}$. This has the advantage of forcing into existence the right-handed $SU(2)_R$ lepton doublet $(\nu, l)_R$, so that ν_R is the Dirac partner of the observed ν_L which belongs to the $SU(2)_L$ lepton doublet $(\nu, l)_L$. At the same time, $B - L$ becomes a gauge symmetry, and its breaking through an $SU(2)_R$ scalar triplet from the 126 of $SO(10)$ also makes ν_R massive, realizing thus the canonical seesaw mechanism for a naturally small Majorana ν_L mass.

Another option [1, 2] is to consider $SO(10) \rightarrow SU(5) \times U(1)_\chi$. This is seldom studied because $SU(5)$ is a grand unified symmetry by itself, so $U(1)_\chi$ is often thought to be unnecessary and uninteresting. Let the breaking of $SU(5)$ to the standard-model (SM) gauge symmetry $SU(3)_C \times SU(2)_L \times U(1)_Y$ be at the scale M_U . If $U(1)_\chi$ survives down to a scale much below M_U and not too far above the electroweak scale, there could be important consequences which have been largely overlooked. In fact, whereas the right-handed neutrino ν_R is a singlet under $SU(5)$, it has a nonzero charge under $U(1)_\chi$. The Higgs doublet which connects u_L to u_R also connects ν_L to ν_R . Hence a Dirac neutrino mass is again obtained and the seesaw mechanism operates as in the left-right case. On the other hand, the detailed phenomenology is very different. Whereas W_R^\pm must exist at the left-right scale M_R , it must be heavier than M_U if $SO(10) \rightarrow SU(5) \times U(1)_\chi$. The Z_χ gauge boson itself has well-defined couplings to the SM particles. Its existence is routinely searched for at the Large Hadron Collider (LHC), with the present mass limit [3, 4] of about 4.1 TeV, which may be improved [5].

In this paper, new fermions and scalars transforming under $U(1)_\chi$ are added to the SM to obtain a number desirable features. With the help of a softly broken Z_2 discrete symmetry, naturally light seesaw Dirac neutrinos [6, 7, 8] may be obtained. The resulting Lagrangian conserves both B and L . Further addition of two scalars with $L = -1, -2$ enables the appearance of self-interacting leptonic dark matter [9]. The analog of leptogenesis (through a heavy singlet Majorana fermion which couples to leptons in the seesaw mechanism) is possible using a heavy singlet Majorana fermion which couples to a scalar diquark and

an antiquark, thereby generating the baryon asymmetry of the Universe. A fermion color octet may also be introduced to support a Peccei-Quinn symmetry to obtain an invisible axion for solving the strong CP problem.

2. SEESAW DIRAC NEUTRINOS

The spinorial 16 representation is again chosen for the three families of quarks and leptons and their decompositions shown in Table 1.

The necessary Higgs scalars for fermion masses belong to the 10 representation, as shown in Table 2. New fermions N, N^c belonging to 126*, 126 respectively are added per family, as well as a Higgs doublet from 144 and a singlet from 16. Note that their Q_χ charges are fixed by the $SO(10)$ representations from which they come. It should also be clear that incomplete $SO(10)$ and $SU(5)$ multiplets are considered here (which is the case for all realistic grand unified models). An important Z_2 discrete symmetry is imposed so that ν^c, N, N^c and η are odd, and the other fields are even. Since Φ_1^\dagger transforms exactly like Φ_2 , the linear combination $\Phi = (v_1 \Phi_1^\dagger + v_2 \Phi_2) / \sqrt{v_1^2 + v_2^2}$ is the analog of the standard-model Higgs doublet, where $\langle \phi_{1,2}^0 \rangle = v_{1,2}$. The Z_2 symmetry is respected by all dimension-four terms of the Lagrangian. It will be broken softly by the dimension-three trilinear term $\mu \sigma \Phi^\dagger \eta$ as well as spontaneously by $\langle \eta^0 \rangle = v_3$. The 4×4 neutrino mass matrix spanning (ν, ν^c, N, N^c) is then given by

$$\mathcal{M}_\nu = \begin{pmatrix} 0 & 0 & 0 & f_\eta v_3 \\ 0 & 0 & f_\sigma u & 0 \\ 0 & f_\sigma u & 0 & m_N \\ f_\eta v_3 & 0 & m_N & 0 \end{pmatrix}, \quad (1)$$

where $u = \langle \sigma \rangle$ which breaks $U(1)_\chi$. The above mass matrix generates a seesaw Dirac neutrino with $m_\nu = f_\eta f_\sigma v_3 u / m_N$, which is naturally small.

3. SCALAR SECTOR

The scalar potential consisting of Φ, η , and σ is given by

$$\begin{aligned} V &= \mu_\Phi^2 \Phi^\dagger \Phi + \mu_\eta^2 \eta^\dagger \eta + \mu_\sigma^2 \sigma^* \sigma + [\mu \sigma \Phi^\dagger \eta + H.c.] \\ &+ \frac{1}{2} \lambda_\Phi (\Phi^\dagger \Phi)^2 + \frac{1}{2} \lambda_\eta (\eta^\dagger \eta)^2 + \frac{1}{2} \lambda_\sigma (\sigma^* \sigma)^2 \\ &+ \lambda_{\Phi\eta} (\Phi^\dagger \Phi) (\eta^\dagger \eta) + \lambda_{\Phi\sigma} (\Phi^\dagger \Phi) (\sigma^* \sigma) \end{aligned} \quad (1)$$

fermion	SO(10)	SU(5)	SU(3) _C	SU(2) _L	U(1) _Y	U(1) _χ	Z ₂
d^c	16	5*	3*	1	1/3	3	+
(ν, e)	16	5*	1	2	-1/2	3	+
(u, d)	16	10	3	2	1/6	-1	+
u^c	16	10	3*	1	-2/3	-1	+
e^c	16	10	1	1	1	-1	+
ν^c	16	1	1	1	0	-5	-
N	126*	1	1	1	0	10	-
N^c	126	1	1	1	0	-10	-

TABLE 1: Fermion content of model.

scalar	SO(10)	SU(5)	SU(3) _C	SU(2) _L	U(1) _Y	U(1) _χ	Z ₂
(ϕ_1^0, ϕ_1^-)	10	5*	1	2	-1/2	-2	+
(ϕ_2^+, ϕ_2^0)	10	5	1	2	1/2	2	+
(η^+, η^0)	144	5	1	2	1/2	7	-
σ	16	1	1	1	0	-5	+

TABLE 2: Scalar content of model.

$$+ \lambda_{\eta\sigma}(\eta^\dagger\eta)(\sigma^*\sigma). \quad (2)$$

The minimum of V satisfies the conditions

$$0 = \mu_\Phi^2 + \lambda_\Phi v^2 + \lambda_{\Phi\eta} v_3^2 + \lambda_{\Phi\sigma} u^2 + \mu v_3 u / v, \quad (3)$$

$$0 = \mu_\eta^2 + \lambda_\eta v_3^2 + \lambda_{\Phi\eta} v^2 + \lambda_{\eta\sigma} u^2 + \mu v u / v_3, \quad (4)$$

$$0 = \mu_\sigma^2 + \lambda_\sigma u^2 + \lambda_{\Phi\sigma} v^2 + \lambda_{\eta\sigma} v_3^2 + \mu v v_3 / u. \quad (5)$$

Assuming that $u \gg v \gg v_3$, the solutions to the above are

$$u^2 \simeq -\mu_\sigma^2 / \lambda_\sigma, \quad (6)$$

$$v^2 \simeq -(\mu_\Phi^2 + \lambda_{\Phi\sigma} u^2) / \lambda_\Phi, \quad (7)$$

$$v_3 \simeq -\mu v / (\mu_\eta^2 + \lambda_{\eta\sigma} u^2). \quad (8)$$

The 3×3 mass-squared matrix spanning $\sqrt{2}[Im(\phi^0), Im(\eta^0), Im(\sigma)]$ is given by

$$\mathcal{M}_I^2 = -\mu \begin{pmatrix} v_3 u / v & -u & -v_3 \\ -u & v u / v_3 & v \\ -v_3 & v & v_3 v / u \end{pmatrix}, \quad (9)$$

which has two zero eigenvalues and one massive eigenstate with

$$m_{\eta_I}^2 = -\mu \left(\frac{v_3 u}{v} + \frac{v u}{v_3} + \frac{v_3 v}{u} \right) \simeq -\frac{\mu v u}{v_3} \simeq \mu_\eta^2 + \lambda_{\eta\sigma} u^2. \quad (10)$$

The 3×3 mass-squared matrix spanning $\sqrt{2}[Re(\phi^0), Re(\eta^0), Re(\sigma)]$ is given by

$$\mathcal{M}_R^2 = 2 \begin{pmatrix} \lambda_\Phi v^2 & \lambda_{\Phi\eta} v_3 v & \lambda_{\Phi\sigma} v u \\ \lambda_{\Phi\eta} v_3 v & \lambda_\eta v_3^2 & \lambda_{\eta\sigma} v_3 u \\ \lambda_{\Phi\sigma} v u & \lambda_{\eta\sigma} v_3 u & \lambda_\sigma u^2 \end{pmatrix} \quad (11)$$

$$-\mu \begin{pmatrix} v_3 u / v & -u & -v_3 \\ -u & v u / v_3 & -v \\ -v_3 & -v & v_3 v / u \end{pmatrix}, \quad (12)$$

which is approximately diagonal with

$$m_{\phi_R}^2 \simeq 2\lambda_\Phi v^2, \quad m_{\eta_R}^2 \simeq m_{\eta_I}^2, \quad m_{\sigma_R}^2 \simeq 2\lambda_\sigma u^2. \quad (13)$$

Note that the $\phi_R - \sigma_R$ mixing (with the assumption that ϕ_R is much lighter than σ_R) is roughly $\lambda_{\Phi\sigma} v u / \lambda_\sigma u^2$ which is naturally suppressed by v/u . As for η_R , its mass is dominated by $-\mu v u / v_3$, and its mixing with ϕ_R and σ_R is suppressed by v_3/v and v_3/u respectively. This justifies the diagonal approximation assumed here.

As a numerical example, let $u = 10$ TeV, $v_3 = 10$ GeV, then $m_\nu = 0.1$ eV is obtained for $f_\eta = f_\sigma = 0.1$ and $m_N = 10^{13}$ GeV. The soft Z_2 breaking parameter μ is then 6 GeV for $m_{\eta_I} = 1$ TeV.

4. CONSERVED BARYON AND LEPTON NUMBERS WITH SELF-INTERACTING DARK MATTER

Because of the $U(1)_\chi$ assignments of the particle content of this model, the resulting Lagrangian conserves both baryon number B and lepton number L even after the symmetry breaking of $U(1)_\chi$ by u, v, v_3 and $SU(2)_L \times U(1)_Y$ by v, v_3 . As usual, the quarks have $B = 1/3, L = 0$ and the leptons (including the Dirac neutrinos) have $B = 0, L = 1$. This means that if a new particle is added, it may be assigned B and L numbers appropriately, according to its assumed interactions with the known quarks and leptons [10, 11]. These assignments lie outside $U(1)_\chi$, hence Q_χ is now not a marker of dark matter, as in previous studies [1, 2].

For example, consider the scalar singlet $\zeta \sim (1, -10)$ from the $\underline{126}$ of $SO(10)$. It has the allowed Yukawa coupling $\zeta^* \nu^c \nu^c$. In conventional models, ζ is assumed to have a vacuum expectation value, thereby breaking $U(1)_\chi$ and giving ν^c a large Majorana mass, breaking thus also lepton number L . Here it may be assumed instead that L is conserved and ζ has $L = -2$. Note that $U(1)_\chi$ is broken instead by σ which may be assigned $L = 0$, together with $L = \pm 1$ for N, N^c , and $L = 0$ for Φ and η .

With ζ as a scalar dilepton which couples only to the Dirac neutrinos, it is then a simple step to consider a scalar singlet $\rho \sim (1, -5)$ from the $\underline{16}$ of $SO(10)$ with $L = -1$ so that it can be self-interacting dark-matter [12] with ζ as its light mediator, as proposed recently [9]. Since ζ decays only to two neutrinos, it does not disrupt the cosmic microwave background (CMB)

scalar	$SO(10)$	$SU(5)$	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_\chi$	L
ζ	126	1	1	1	0	-10	-2
ρ	16	1	1	1	0	-5	-1

TABLE 3: Leptonic scalars for self-interacting dark matter.

scalar/fermion	$SO(10)$	$SU(5)$	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_\chi$	B
h_1 (scalar)	16*	5	3	1	-1/3	-3	-2/3
h_2 (scalar)	10	5	3	1	-1/3	2	-2/3
ψ (fermion)	45	24	1	1	0	0	1

TABLE 4: New particles for baryogenesis.

scalar/fermion	$SO(10)$	$SU(5)$	$SU(3)_C$	$SU(2)_L$	$U(1)_Y$	$U(1)_\chi$	PQ
Ω (fermion)	45	24	8	1	0	0	1
S (scalar)	54	24	1	1	0	0	2

TABLE 5: New particles for strong CP conservation through the axion.

from its enhanced production at late times due to the Sommerfeld effect. It removes an important objection [13, 14] to models where the light mediator decays to electrons and photons, usually through Higgs mixing, which is forbidden here by L conservation.

5. BARYOGENESIS

Since lepton number is strictly conserved, the usual mechanism of generating the baryon asymmetry of the Universe through leptogenesis is not possible. However, the analog process of having a heavy Majorana fermion ψ decaying to $B = \pm 1$ final states [15] may be implemented with the addition of two scalar diquarks $h_{1,2}$, as shown in Table 4.

The allowed couplings involving the new particles are

$$h_2 u d, h_2^* u^c d^c, h_1 d^c \psi, h_1^* h_2 \sigma. \quad (1)$$

Assuming a large Majorana mass for ψ which breaks B to $(-1)^{3B}$, it may now decay to both $h_1^* d^c$ ($B = 1$) and $h_1 d^c$ ($B = -1$). The subsequent decay of h_1 to $u^c d^c$ through $h_1 - h_2$ mixing from $\langle \sigma \rangle$ establishes a baryon asymmetry in analogy to the case of a lepton asymmetry if ν^c were a heavy Majorana fermion. The one-loop vertex [16] and self-energy [17] diagrams which contribute to the CP asymmetry and thus the B asymmetry are depicted in Fig. 1. They are completely analo-

B asymmetry is converted to a $B - L$ asymmetry through the spherulons, again in analogy to what happens in leptogenesis.

A possible scenario is high-scale baryogenesis with $m_{\psi_1} \sim 10^{13}$ GeV. Instead of three families of leptons in leptogenesis, just one set of scalar diquarks $h_{1,2}$ is needed. The CP asymmetry generated by the decay of ψ_1 assuming that ψ_2 is much heavier is given by

$$\epsilon = -\frac{3}{16\pi} \frac{m_{\psi_1}}{m_{\psi_2}} \frac{\text{Im}[(f_1 f_2^*)^2]}{|f_1|^2}, \quad (2)$$

where the ψ_1 decay rate is $\Gamma_1 = |f_1|^2 m_{\psi_1} / 8\pi$. Consider the parameter $K = \Gamma_1 / H(T = m_{\psi_1})$, where the Hubble parameter is $H = 1.66 \sqrt{g_*} (T^2 / M_{Pl})$, as a measure of the deviation from equilibrium. If $K \ll 1$, which means $|f_1| \ll 0.02$, then the baryon asymmetry is of order ϵ / g_* . Setting this to 10^{-10} , and assuming $m_{\psi_2} / m_{\psi_1} = 6$, then $|f_2| = 10^{-3}$ if the relative phase between f_1 and f_2 is of order 1.

6. AXIONIC DARK MATTER

To obtain an axionic solution to the strong CP problem, a colored fermion is needed which has an anomalous Peccei-Quinn charge. Instead of the usual quark triplet, a fermion color octet, such as the gluino of supersymmetry, may be used [18]. In a nonsupersymmetric context, it may just be any fermion color octet [19] unrelated to the gluon. Here it is

called Ω and it obtains a large Majorana mass through the interaction $S^* \Omega \Omega$, so that the dynamical phase of S becomes the invisible axion which is a component of dark matter.

7. CONCLUDING REMARKS

In the context of $U(1)_\chi$, new light is shed on some of the outstanding problems in particle physics and astroparticle physics. It is shown how naturally light Dirac neutrinos may be obtained in a seesaw mechanism which is usually reserved for considering Majorana neutrinos. With light Dirac neutrinos, an elegant solution to an important problem in self-interacting dark matter may also be solved. The light scalar mediator here is a dilepton and decays only to two neutrinos, so it does not

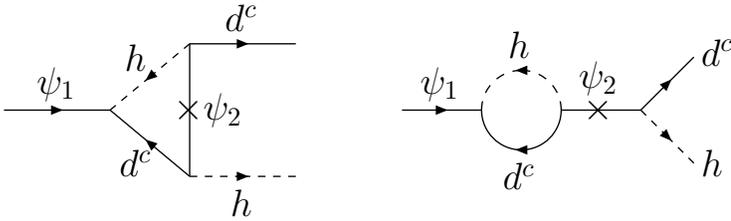


FIGURE 1: One-loop diagrams for baryogenesis.

gous to those of leptogenesis where $\psi_{1,2}$ are replaced by $\nu_{1,2}^c$, the scalar diquark h by ϕ^+ , and d^c by l . However, since the Yukawa interactions $h d^c \phi_{1,2}$ are not constrained by lepton masses as in $\phi^+ l \nu_{1,2}^c$, the desirable asymmetry is easily obtained. Note of course that $m_h \ll m_{\psi_1} < m_{\psi_2}$ is assumed. The resulting

disrupt the cosmic microwave background at late times. With the conservation of lepton number, the possibility of breaking baryon number B to baryon parity, i.e. $(-1)^{3B}$, allows baryogenesis to occur, in analogy to leptogenesis, from the decay of a heavy Majorana fermion carrying $B = 1$. To explain strong CP conservation, a Majorana fermion color octet with anomalous Peccei-Quinn charge is postulated. It acquires a large mass through its coupling to a singlet scalar, the dynamical phase of which becomes the invisible axion and contributes to dark matter.

8. ACKNOWLEDGEMENT

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References

- [1] E. Ma, Phys. Rev. **D98**, 091701(R) (2018) [arXiv:1809.03974].
- [2] E. Ma, LHEP (Special Issue); arXiv:1810.06506 [hep-ph].
- [3] ATLAS Collaboration, M. Aaboud *et al.*, JHEP **1710**, 182 (2017) [arXiv:1707.02424].
- [4] CMS Collaboration, A. M. Sirunyan, A. Tumasyan *et al.*, JHEP **1806**, 120 (2018) [arXiv:1803.06292].
- [5] S. J. D. King, S. F. King, and S. Moretti, Phys. Rev. **D97**, 115027 (2018) [arXiv:1712.01279].
- [6] P. Roy and O. U. Sanker, Phys. Rev. Lett. **52**, 713 (1984).
- [7] E. Ma, Phys. Rev. Lett. **89**, 041801 (2002) [hep-ph/0201083].
- [8] E. Ma, N. Pollard, R. Srivastava, and M. Zakeri, Phys. Lett. **B750**, 135 (2015) [arXiv:1507.03943].
- [9] E. Ma, Mod. Phys. Lett. **A33**, 1850226 (2018) [arXiv:1805.03295].
- [10] E. Ma, Phys. Rev. Lett. **115**, 011801 (2015) [arXiv:1502.02200].
- [11] E. Ma, Mod. Phys. Lett. **A32**, 1730007 (2017) [arXiv:1702.03281].
- [12] A. Kamada, M. Kaplinghat, A. B. Pace, and H.-B. Yu, Phys. Rev. Lett. **119**, 111102 (2017) [arXiv:1611.02716].
- [13] S. Galli, F. Iocco, G. Bertone, and A. Melchiorri, Phys. Rev. **D80**, 023505 (2009) [arXiv:0905.0003].
- [14] T. Bringmann, F. Kahlhoefer, K. Schmidt-Hoberg, and P. Walia, Phys. Rev. Lett. **118**, 141802 (2017) [arXiv:1612.00845].
- [15] E. Ma, Phys. Rev. Lett. **60**, 1363 (1988).
- [16] M. Fukugita and T. Yanagida, Phys. Lett. **B174**, 45 (1986).
- [17] M. Flanz, E. A. Paschos, and U. Sarkar, Phys. Lett. **B345**, 248 (1995) [hep-ph/9411366].
- [18] D. A. Demir and E. Ma, Phys. Rev. **D62**, 111901(R) (2000) [hep-ph/0004148].
- [19] E. Ma, T. Ohata, and K. Tsumura, Phys. Rev. **D96**, 075039 (2017) [arXiv:1708.03076].